GPU Beamforming for Phased Array Feeds

Justin (Jay) Smallwood Industry PhD (iPhD) Candidate Swinburne University / Fourier Space / CSIRO

11th April, 2025

Outline

- ▶ Who am I?
- What is a Phased Array Feed (PAF), & What's a Beam and How Do We Form Them?
- ▶ Why GPUs for Beamforming?
- Science Opportunities!
- ► What I've Done & What's Next?

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➤ 2011 - 2017 University of Melbourne BCom (Economics), MSci (Mathematics & Statistics). GPU Beamforming for Phased Array Feeds

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➤ 2018 - 2025
Goldman Sachs & Catalyst Funds Management
Equity Trader, Strategist, & Lead Developer.

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 iPhD Candidate.

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Oh...ATNF!

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Who Am I Working With?

▶ Swinburne: Professors Adam Deller & Matthew Bailes,

CSIRO: Dr Chris Phillips,

Fourier Space: Andrew Jameson.

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Phased Array Feeds (PAF)

Made up of 188 receivers in a checkerboard pattern.

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Phased Array Feeds (PAF)

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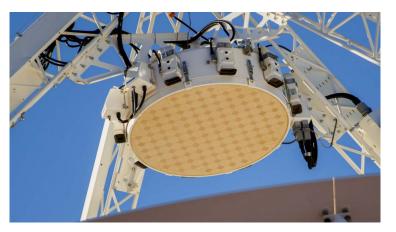


Figure: A PAF on an ASKAP antenna. Credit: CSIRO.

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Phased Array Feeds

- ► Can be mathematically steered to create many beams on the sky simultaneously.
- ► E.g. ASKAP creates 36 dual-polarization beams on the sky for a field-of-view of 30 square degrees.
- ► A CryoPAF is being commissioned at Murriyang.

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Beamforming

How do we combine all these antennas into a beam?

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ow It's Going

How do we combine all these antennas into a beam? **Ingredients:**

- 188 separate data streams from individual antenna elements,
- Compute Power.

Method:

- Weight & sum the data streams to form beams.
- At its simplest, weights are delays between antennas to maximize sensitivity to signals from a particular direction.
- Different weighting schemes can be attractive as well (shaped beamforming).

Beamforming

- \triangleright Express these weights as a vector ω_i ,
- \triangleright Stack beam weights into a matrix ω ,
- Matrix multiply weights against data to create many beams in parallel:

Beams = $\omega \times \mathsf{Data}$.

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How is beamforming currently done?

Beamforming uses Field Programmable Gate Arrays (FPGAs).

- Pros:
 - ► FPGAs are extremely flexible,
 - Excels at base signal processing,
 - Can often hit peak performance as you can tailor architecture to your problem.

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Beamforming uses Field Programmable Gate Arrays (FPGAs).

- Pros:
 - ► FPGAs are extremely flexible,
 - Excels at base signal processing,
 - Can often hit peak performance as you can tailor architecture to your problem.
- Cons:
 - Increased development time,
 - Difficult to iterate and prototype on,
 - Difficult to do more than base signal processing,
 - Limited pool of talent, very complex to develop & debug!

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Why GPUs?

- ► Faster development cycle,
- Built for parallelism,
- Consumer availability,
- Software-based approach rather than hardware.
 - ► Can use C / C++ rather than Verlog / VHDL,
- Larger pool of talent.

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Comparison of Different Methods

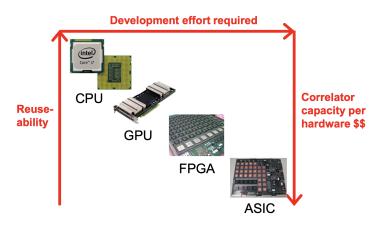


Figure: Credit: A Deller.

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Current GPU-based Beamformers

- ► COBALT LOFAR (Broekema et al. 2018).
- CHIME (Ng et al. 2017).
- ► BLADE Allen Telescope Array.

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Science Opportunities

PhD aim: To demonstrate the flexibility of Phased Array Feeds and GPU-based beamformers for prototyping & developing algorithms to solve observational problems.

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Science Opportunities

PhD aim: To demonstrate the flexibility of Phased Array Feeds and GPU-based beamformers for prototyping & developing algorithms to solve observational problems.

Two Main Areas of Application:

- Radio Frequency Interference Mitigation,
- Real-time Identification of Astronomical Signals.

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Radio Frequency Interference Mitigation

A multi-pixel receiver like a PAF has a significant advantage when detecting RFI:

Celestial source will be in some of the beams, but RFI will be in all the beams! GPU Beamforming for Phased Array Feeds

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Radio Frequency Interference Mitigation

A multi-pixel receiver like a PAF has a significant advantage when detecting RFI:

- Celestial source will be in some of the beams, but RFI will be in all the beams!
- Can use methods like spatial filtering to subtract off common signals between beams.
- Requires correlation matrix between different beams highly parallelizable.

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Shaped Beamforming

- ► (Kocz et al. 2010, 2012) explores a spatial filtering approach to removing RFI.
 - In the beamforming phase, can weight & sum in particular ways to get null response towards a particular direction.
 - Requires calculation of array correlation matrix (ACM).
 - ► GPU-based beamforming may give sufficient performance to allow this in real-time.
- ► (Finlay et al. 2023) provides a method to deal with interference sources moving on a fixed trajectory.
 - ► Further work suggested using GPUs to correlate across frequency bands as well as beams.

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Real-time Identification of Astronomical Signals

- ► Can we detect in real-time if there is something interesting in a beam? I.e. Fast Radio Burst.
- Use of machine learning models to flag interesting data points.
- GPU-based system allows rapid development & feedback loop.
- ► GPUs allow easy integration with machine learning libraries.

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Learning astronomy basics, signal processing, interferometry. GPU Beamforming for Phased Array Feeds

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- Learning astronomy basics, signal processing, interferometry.
- Created a Python package for decoding CODIF data, pycodif (https://github.com/jdgsmallwood/pycodif).

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- ► Leveraged sample jimble data from the Murriyang CryoPAF commissioning to write a crude beamformer in Python.

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- Leveraged sample jimble data from the Murriyang CryoPAF commissioning to write a crude beamformer in Python.
- Created CUDA kernels for an initial (naive) beamforming implementation with 1 beam.
- Side Project: Satellite Two Line Elements (TLEs), adjusting a TLE to agree with new observed positions.

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Extending CUDA kernels for formation of many beams.

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- Extending CUDA kernels for formation of many beams.
- Prototyping optimized kernels using cuBLAS & tensor-core operations.

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- Extending CUDA kernels for formation of many beams.
- Prototyping optimized kernels using cuBLAS & tensor-core operations.
- Exploring implementations for RFI-mitigation techniques for stationary sources (spatial filtering, etc.).

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- Extending CUDA kernels for formation of many beams.
- Prototyping optimized kernels using cuBLAS & tensor-core operations.
- Exploring implementations for RFI-mitigation techniques for stationary sources (spatial filtering, etc.).
- Extending to moving sources (satellites, aircraft).

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